

Laser Plasma Interaction

Mohammed Shihab



Physics Department, Faculty of Science Tanta University



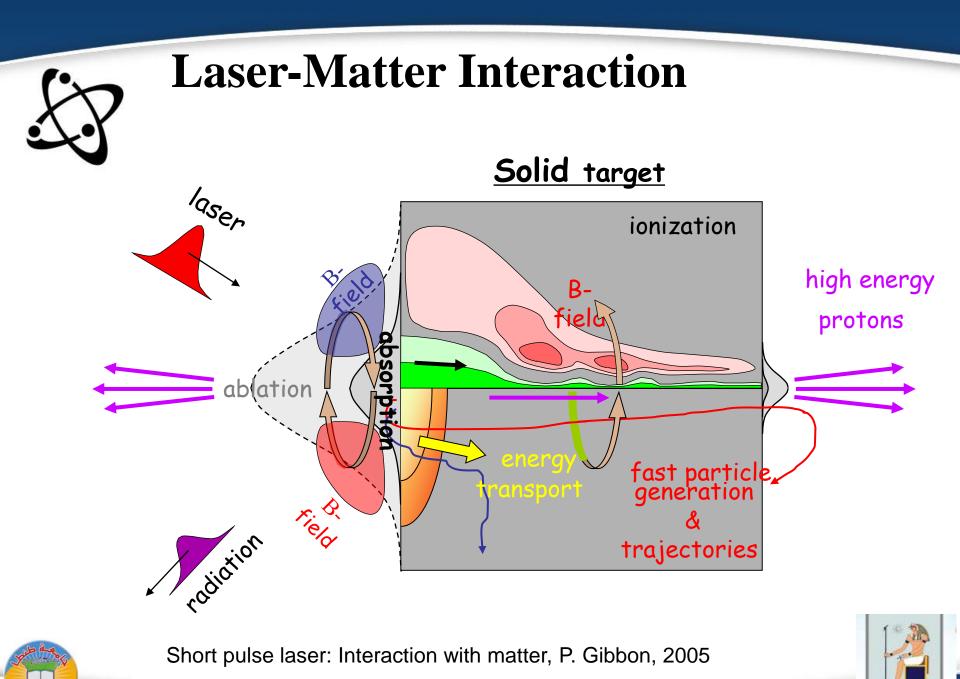


Outline

- Introduction.
 - What is plasma?
 - What is Laser?
 - Laser Ionization dynamics.
 - ns versus fs lasers (Plasma parameters).
- Applications
 - Laser induced breakdown spectroscopy (LIBS).
 - Plasma deposition and etching.
 - UV and x-ray production.
 - Warm dense matter.





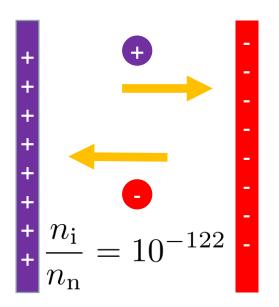


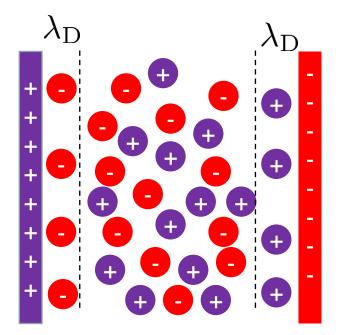
ية العنوم امعة طلطا



What is plasma?

<u>The plasma</u> is a quasineutral gas of charged and neutral particles which exhibits collective behavior.







PLASMA Somthing molded or fabricated





What is Laser?

<u>The laser is a beam of photons characterized with :</u> Monochromatic, Directionality, Coherent, High intensity.

- Conventional Lasers:
 - Active medium
 - Pumping
 - Population inversion
 - Stimulated emmission
 - Resonator



Wicked Lasers



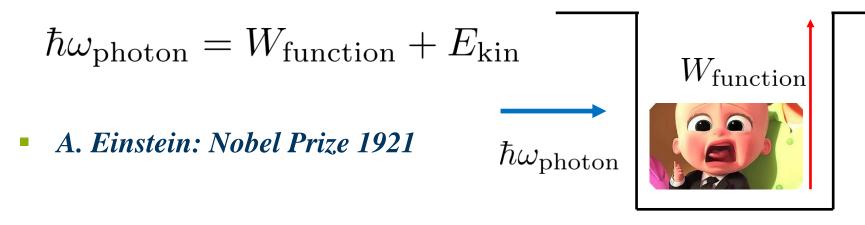
X-ray Free electron Lasers !!!!

Pater Page



Light matter interaction

 <u>Photoelectric effect:</u> Ionization takes place if the photon energy greater than the work function of the matter.



- Ti:Sa Laser 800 nm with photon energy $\hbar\omega=1.5{
m eV}$

Material	Cs	Li	Na	Hg	Au
W. function	2.1 eV	2.9 eV	2.3 eV	4.5 eV	5.5

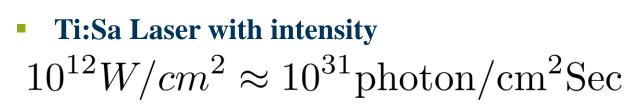


Laser matter interaction I

- <u>Multiple Photon ionization</u> : one single atom can interact with multiple photons at the same time.
- The rate of ionization depends on the laser intensity

$$\Gamma^n = \sigma_{\rm n} I_{\rm L}^{\rm n}$$

 $n\hbar\omega$



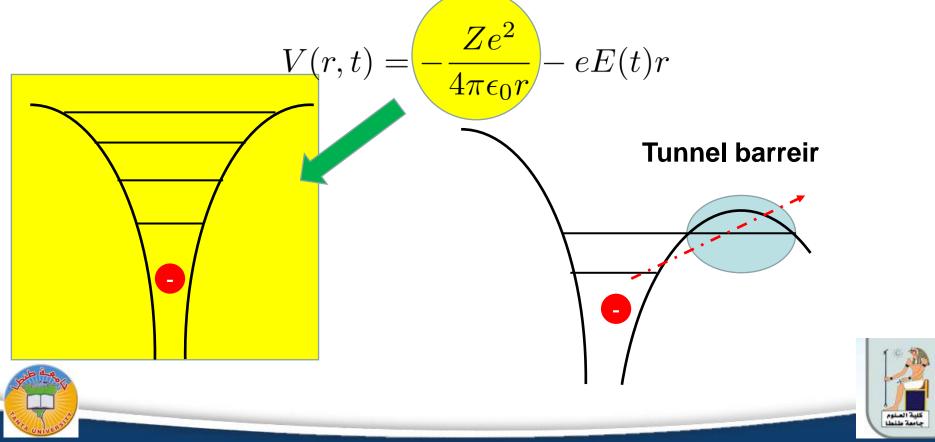
Material	Ionization energy	N- photons
H ₂	13.6 eV	9 photons
Не	24.5 eV	16 photons





Laser matter interaction II

<u>Tunnel ionization</u>: The potential of the laser field can modify the Coulomb potential of an electron in an atom and forms a potential barrier. It depends on the barrier width and height.





Laser matter interaction III

Keldysh Parameter: depends on the ionization energy of an electron ξ to the pondermotive potential $e\Phi$.

$$\gamma = \sqrt{\xi/e\Phi} = \sqrt{\xi/(\lambda_L^2 I_L)}$$

Tunnel ionization $\gamma \leq 1$

 $\mathbf{MPI} \quad \gamma > 1$

For Ti:Sa Laser interacts with He:

Intensity	Keldysh parameter	Ionization
10^12 w/cm^2	14	MPI
2 * 10^14 W/cm^2	1	Tunnel

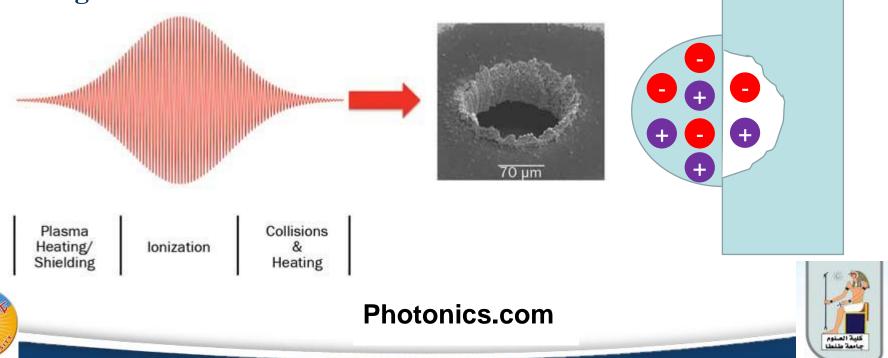
Barrier suppression ionization (BSI): At higher laser intensityies electons can leave atoms without tunneling : dominant above 10^18 W/cm^2.

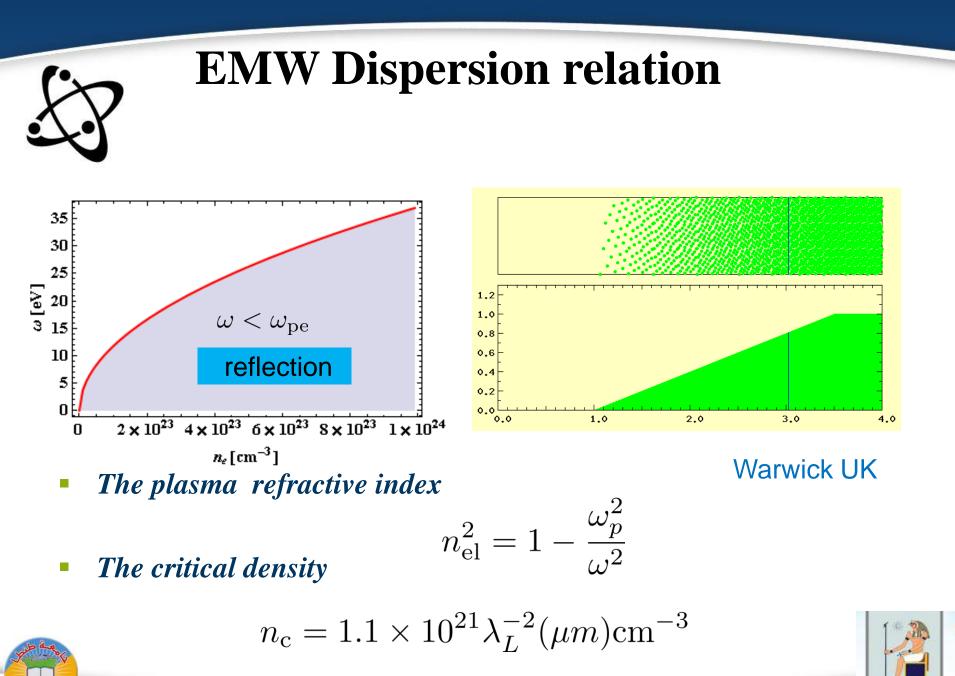


£?

Plasma creation over solid targets

- The first seed of electrons are generated via MPI, Tunnel ionization, or , BSI.
- The free electrons gain energy from the laser electric field, then make further ionization by collision with neutral particles in the target.







Nano-second Pulses I

 The target is heated via electron-ion collisions to 10s to 100s eV depends on the laser intensity

$$T_{\rm e} \approx 13 \times 10^7 (I_{\rm abs} (W/{\rm cm}^2)/n_{\rm e})^{2/3} eV$$

• The plasma pressure created during heating causes ion blow-off (ablation) at the sound speed

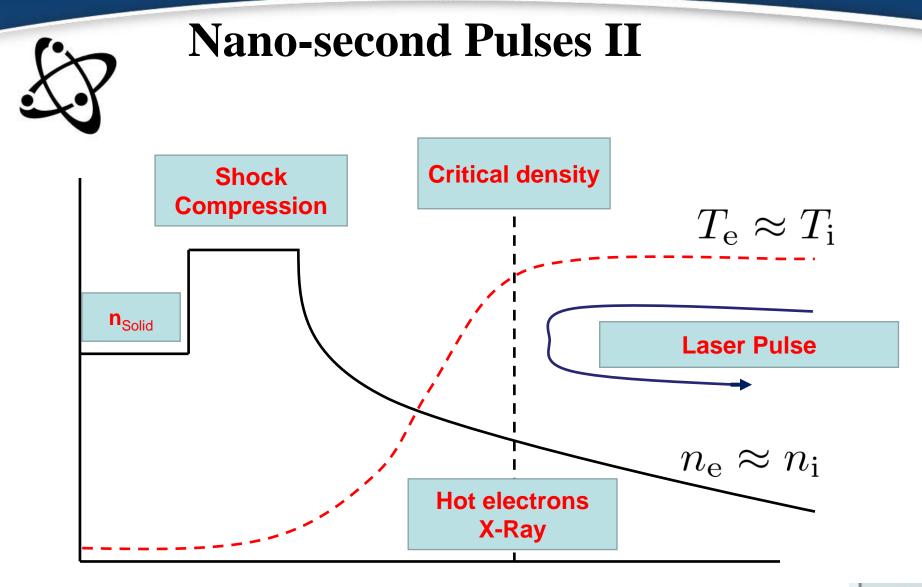
$$C_{\rm s} \approx 3.1 \times 10^7 (T_{\rm KeV} Z^*)^{1/2} {\rm cm/s}$$

 Because of ablation, density profile decreases exponentially with a scale length of

$$L = C_{\rm s} \tau_{\rm L}$$











Absorption Mechanisms I

 Inverse Bremsstrahlung absorption: The electrons, while oscillating under the action of the laser electric field, collide with the ions giving rise to transfer of electromagnetic energy to the plasma.

$$K_{\rm ib} \approx 3.1 \times 10^{-7} Z n_{\rm e}^2 \ln(\Lambda) \omega_L^{-2} [1 - \frac{\omega_p^2}{\omega_L^2}]^{-1/2} T_{\rm eV}^{-3/2} {\rm cm}^{-1}$$

• The fraction of absorbed laser energy after a propagation over a distance L in a uniform plasma is

$$\alpha_{\rm abs} = 1 - exp[-K_{\rm ib}L]$$

 Ion turbulence absorption: Inverse Brem. Absorption increases significantly when ion motion is correlated.



Absorption Mechanisms II

- Resonance absorption:
 - Laser radiation obliquely incident on a plasma and with $(\vec{E}_L \cdot \vec{\nabla}(n_{\rm e}) \neq 0)$ can excite resonant longitudinal plasma oscillations at critical density surface.
 - The damping of the excited electron waves leads to conversion of electromagnetic laser energy into thermal energy.
- The amplitude of the electrons oscillations depends on the laser polarization: $\vec{F} = \vec{\nabla}(m_{e})$

$$\delta n_{\rm e} \propto \frac{E_L \cdot \nabla(n_{\rm e})}{n_{\rm e} - n_{\rm c}}$$

• S-Polarized $(\vec{E}_L \cdot \vec{\nabla}(n_e) = 0)$ can not excite plasmons or Langmuir waves.



Absorption Mechanisms III

Parametric instabilities: Wave-wave interaction

$$\omega_0 = \omega_1 + \omega_2 \quad K_0 = K_1 + K_2$$

Photon
 Photon + Acoustic
 Stimulated Brillouin scattering

 Photon
 Photon + Plasmon
 Stimulated Raman scattering

 Photon
 Acoustic + Plasmon
 Decay instability

Photon — Plasmon + Plasmon

Tow-Plasmon instability





Femto-second Lasers

- Brunel effect: Intense laser pulses incident on sharp overdense plasmas, pulls electrons into the vacuum and then back them into the plasma.
- Filamentation: Energetic beam of electron penetrate into the core of the target producing ionization.
- Surface wave: Sharp overdense plasma boundary supports surface waves; electron fluctuations accompanied with electromagnetic waves.

• And more



Radiation Hydrodynamic

- Governing equations in planar geometry:
- The independent Lagrangian mass variable: The spatial grids moves with the fluid $dm_0 = \rho(r)dr$
- The momentum conservation equation $\frac{\partial u}{\partial t} = -\frac{\partial}{\partial t}(P+q)$

•
$$P$$
 is the total pressure due to electrons, ions, and radiation,
 q is Neumann artificial viscosity

Conservation of energy

$$C_{\rm v,e} \frac{\partial T_{\rm e}}{\partial t} = \frac{\partial}{\partial m_0} \xi_{\rm e} \frac{T_{\rm e}}{\partial r} - \omega_{\rm ei} (T_{\rm e} - T_{\rm i}) + R_{\rm Abs} - R_{\rm Emis} + S_{\rm e} - \dots$$
$$C_{\rm v,i} \frac{\partial T_{\rm i}}{\partial t} = \frac{\partial}{\partial m_0} \xi_{\rm e} \frac{T_{\rm i}}{\partial r} + \omega_{\rm ei} (T_{\rm e} - T_{\rm i}) - \dots$$





£P

Local thermodynamic equilibrium

- Thermodynamic equilibrium: require that plasma fulfils four energy distributions; radiative, kinetic, excitation, and ionization energy, termed, the Planck, Maxwell, Boltzmann, and Saha relations; each energy exchange process must be balanced by its inverse. (Not found in man made plasmas).
- Local thermodynamic equilibrium; when collisional processes are more important that radiative, one can find a temperature based on Maxwell, Boltzmann, and Saha distributions.
- All atomic processes must be considered: radiative ionization, radiative excitation, radiative recombination, impact ionization, dielectronic recombination, three body recombination,





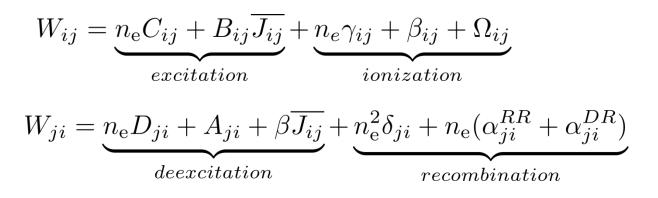
£P

Atomic physics calculations

Atomic rate equations:

$$\frac{dn_i}{dt} = -n_i \sum_{i \neq j}^{N_L} W_{ij} + \sum_{i \neq j}^{N_L} n_j W_{ji}$$

Upward (depopulating) and downward (populating) transitions:

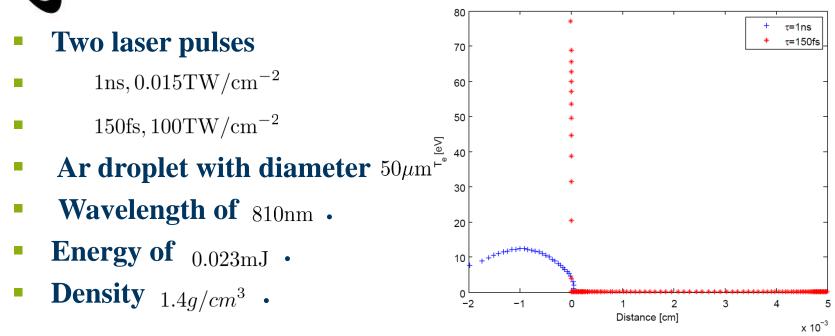


Continuum lowering effects are considered





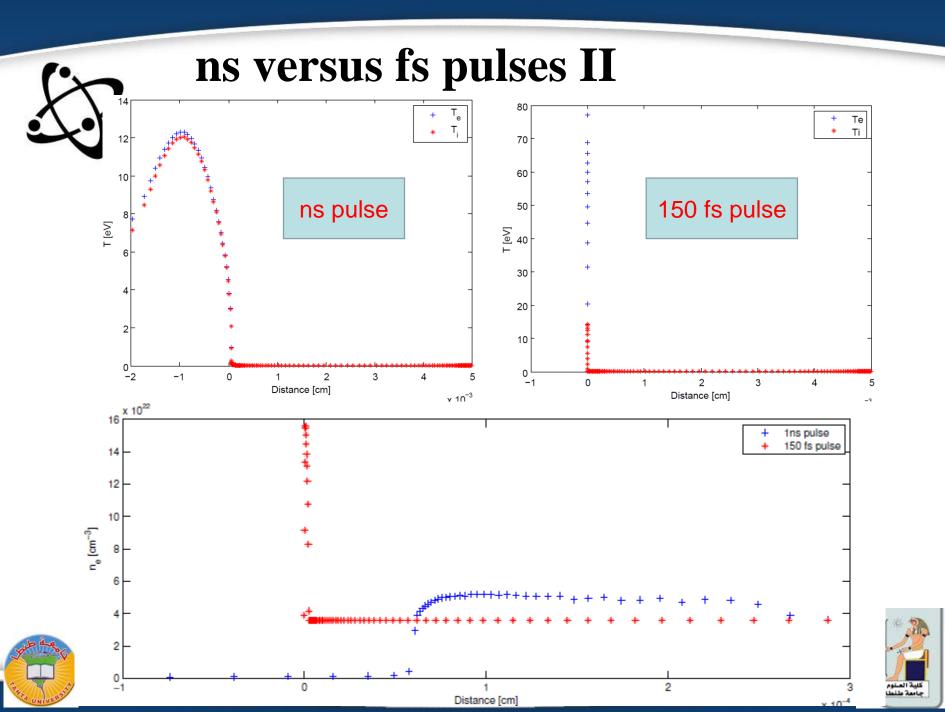
ns versus fs pulses I



 Electron-ion collision and ion hydrodynamic expansion time are in sub ps.









Laser Droplet interaction movie

Entry #: V0016 Laser impact on a drop

Alexander L. Klein¹, Claas Willem Visser¹, Wilco Bouwhuis¹, Henri Lhuissier², Chao Sun¹, Jacco H. Snoeijer¹, Emmanuel Villermaux³, Detlef Lohse¹ and Hanneke Gelderblom¹

¹ Physics of Fluids Group, Faculty of Science and Technology, University of Twente, The Netherlands ² Laboratoire Matière et Systèmes Complexes, Université Paris Diderot, France

³ IRPHE, Aix-Marseille Université, France



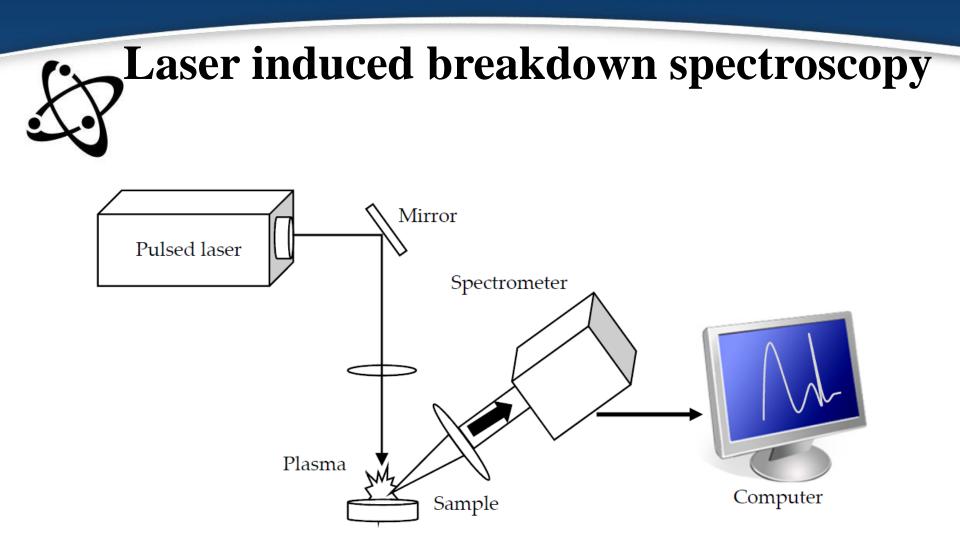




Applications





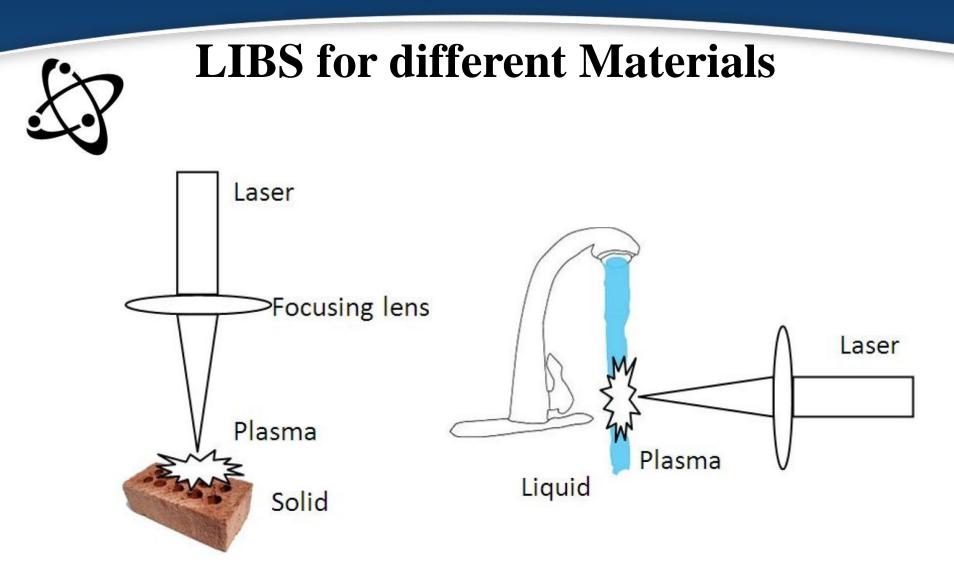


The conventional LIBS configuration









Solid and liquid samples

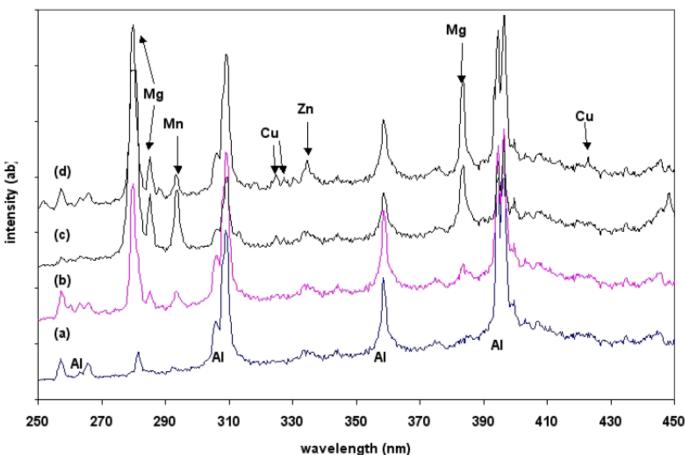


InTech





Al alloys spectrum



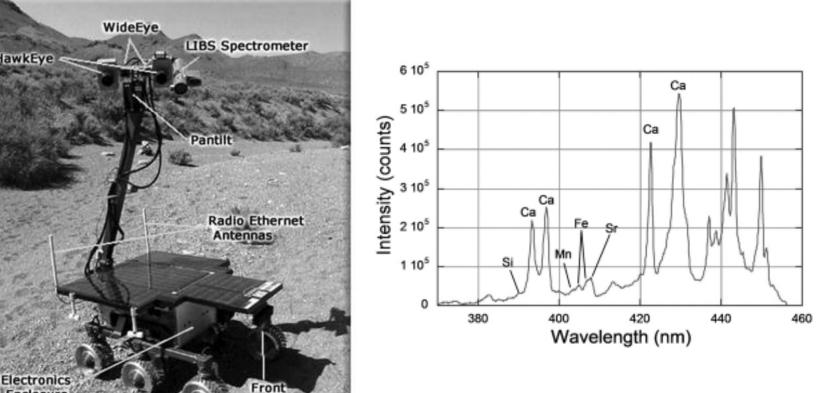
LIBS for Al alloys (a) Pure Al, (b) 3003 alloy, (c) 2024-T3 alloy, and (d) 7057-T6 alloy.
 InTech

للية المنوم مامعة طلطا



Enclosure

Study your samples in the field!!

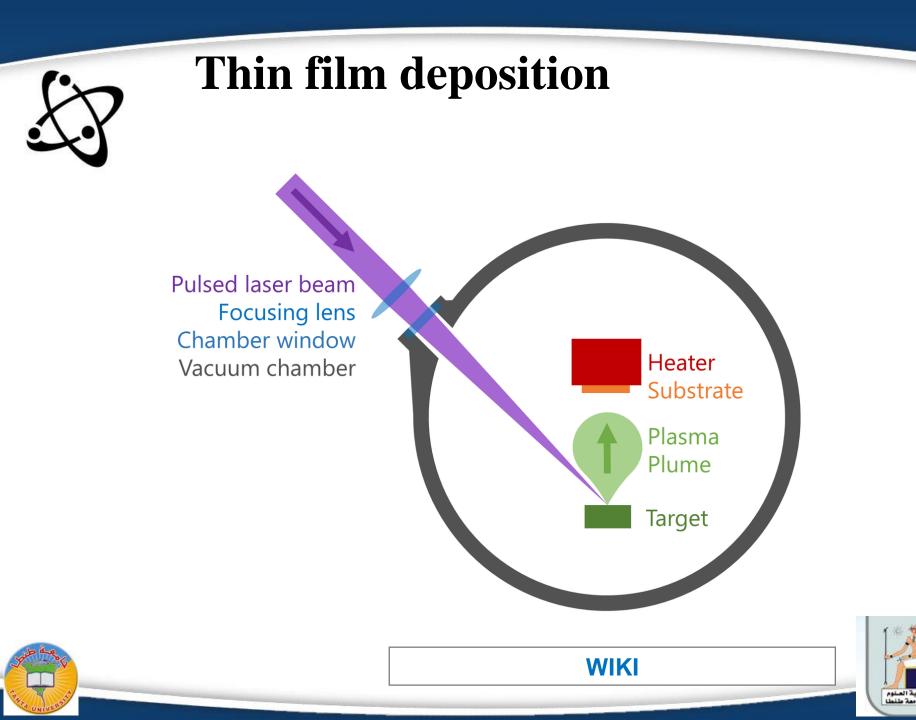


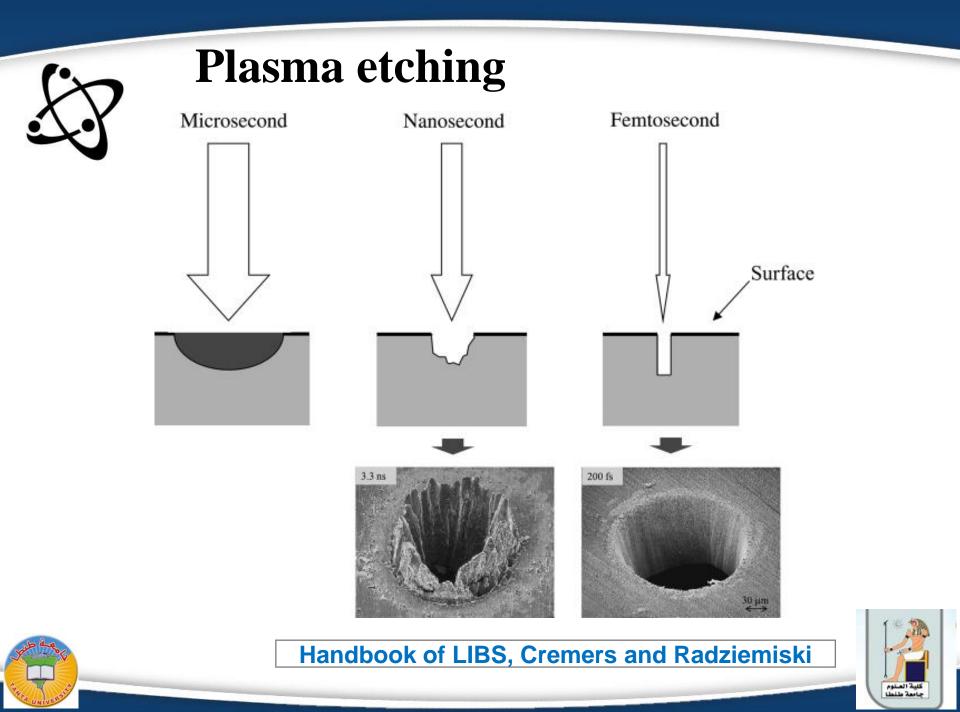
• The k9 rover in the field (Nevada, USA).

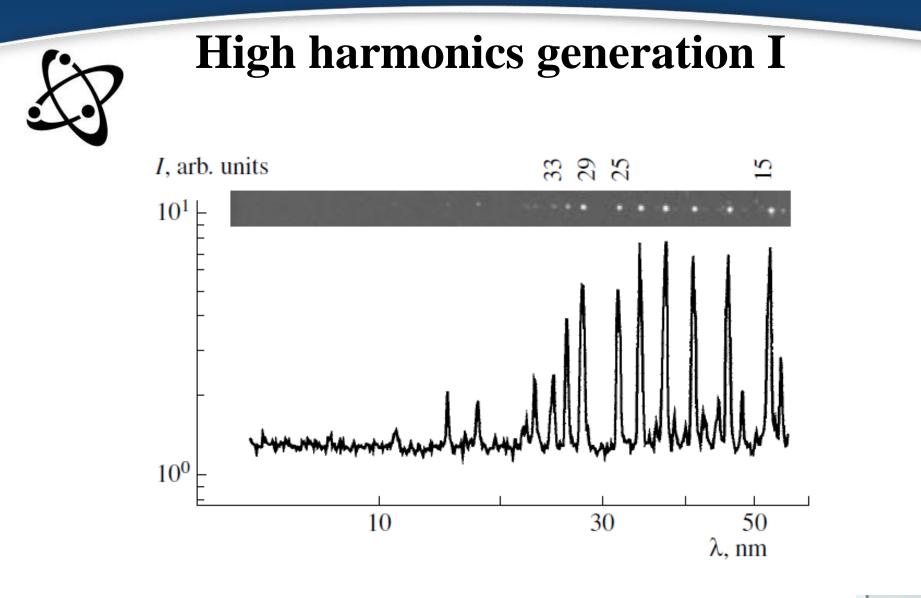
Hazcams









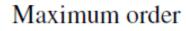


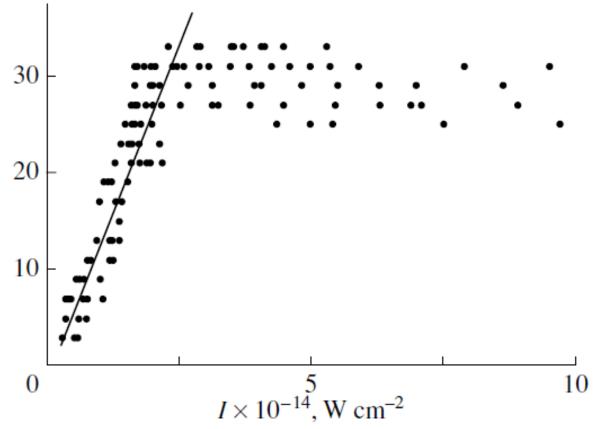
Ganeev and Kuroda, optics and spectroscopy 2006





High harmonics generation I





Ganeev and Kuroda, optics and spectroscopy 2006



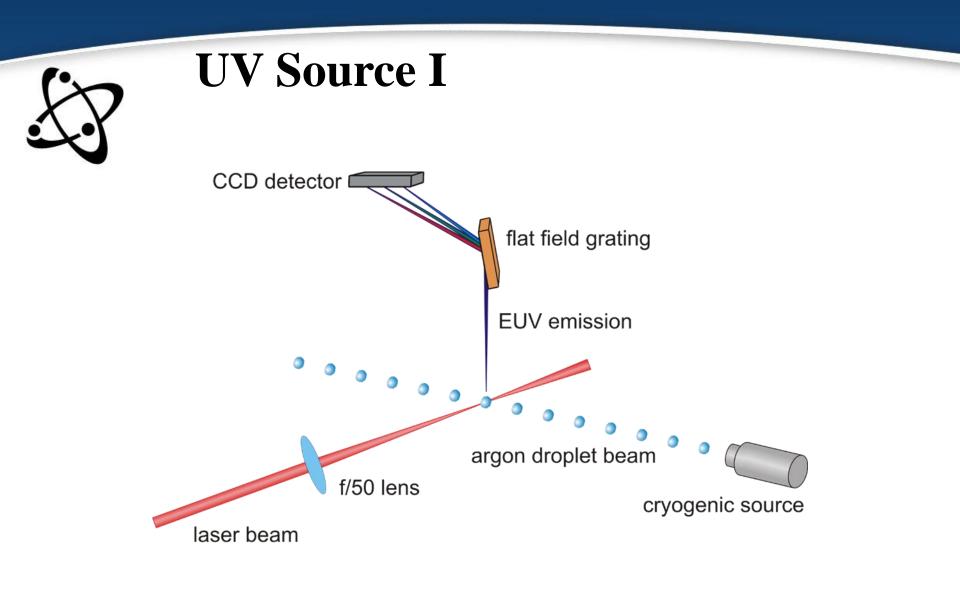
X-ray Production

- Laser produced hot plasmas which can produce x-ray.
- Three mechanisms are exist:
 - Bresmsstahlung or (Free-Free emission)
 - Recombination radiation (Free-bound emission).
 - Line radiation (bound-bound emission).
- The laser to x-ray conversion efficiency E_x/E_L depends on the target atomic number, the laser intensity, the laser wavelength;

$$E_x/E_L = 6.3 \times 10^{-10} Z \lambda_{\mu m}^{-0.48} \left(\frac{I_0}{1+I_0^2}\right)^{0.46}$$

H.C. Pant, plasma physics 1992



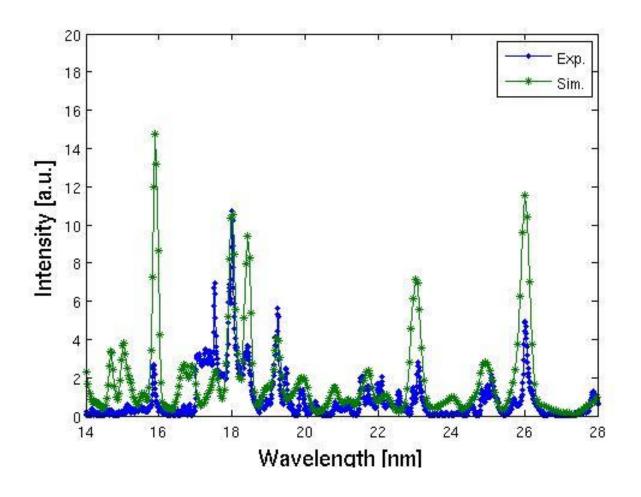








UV Source II

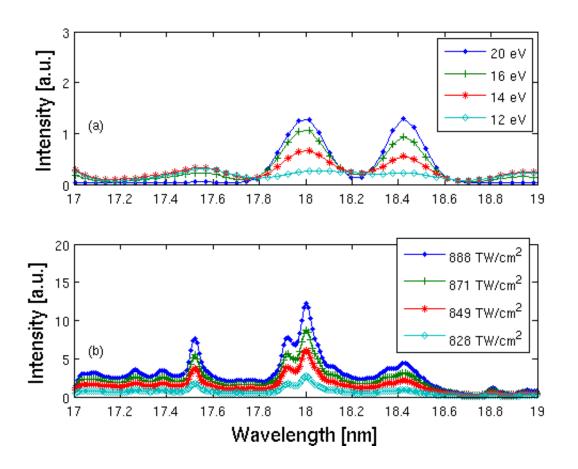


R. Irsig, M. Shihab et al 2018





UV Source III



R. Irsig, M. Shihab et al 2018

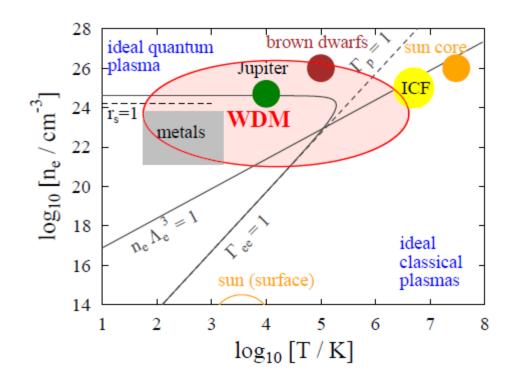








Warm Dense Matter



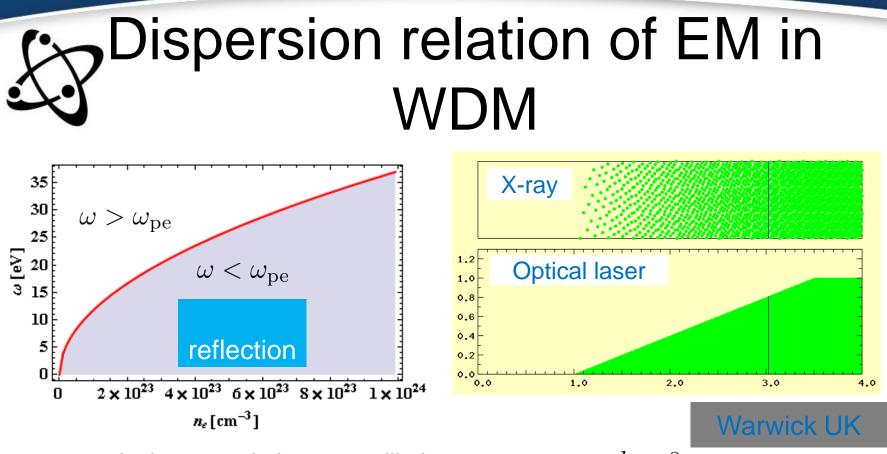
K. Wünsch

- Temperature of few electronvolts
- Solid state density and beyond
- ICF, shock experiments, giant planets, and brown dwarfs
- Theories of solid, condensed matter, or ideal plasma are not valid
- No single theoretical model describes the behavior of WDM
 - Partial ionization
 - Arbitrary degeneracy
 - Strong ionic correlations



Glenzer et al PRL 98 065002(2007)





• At the natural plasma oscillation: $\omega_{
m pe} = \omega
ightarrow k = 0$

• At the cut off, the wave is reflected:

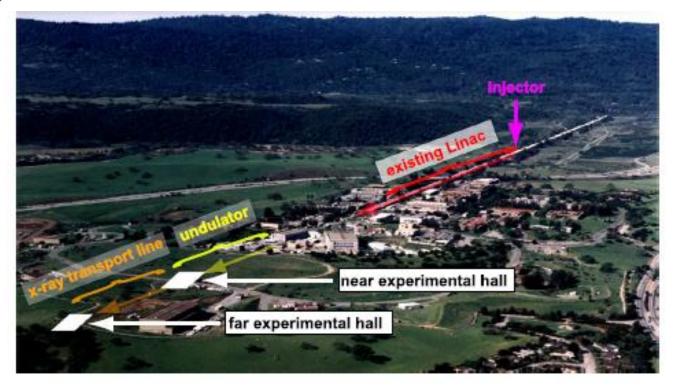
$$\omega_{\rm pe} > \omega \to k = i\kappa$$

WDM is transparent in x-ray regime:

$$n_{\rm e} = 10^{24} {\rm cm}^{-3} \rightarrow \lambda \le 33 nm$$



Set-up of an XRTS experiment



S.H. Glenzer et al., (2016): Stanford University





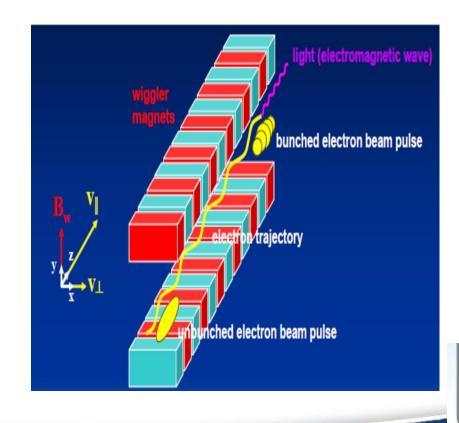
Free electron laser

✤The free electron laser (FEL) is a device that transforms the kinetic energy of a relativistic electron beam into electromagnetic (EM) radiation.

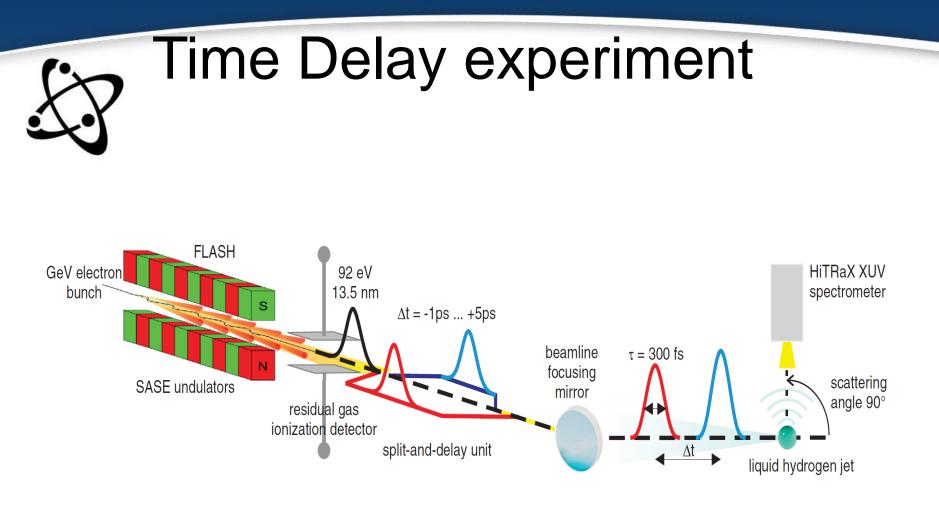
✦Electrons in an FEL are not

bound to atoms or molecules.

The "free" electrons traverse a series of alternating magnets, called a "wiggler," and radiate light at wavelengths depending on electrons' energy, wiggler period and magnetic field.







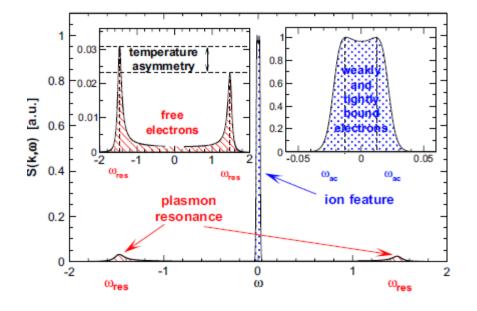
U. Zastrau 2014: FLASH(Hamburg)







XRTS features



A. Höll et al., HEDP 3, 120(2007)

- Thomson scattering has two distinct features:
 - Inelastic scattering (frequency shifted) from free electrons and bound free transitions
 - Unshifted Rayleigh peak (elastic) due to electrons comoving with the ions
- The electrons in partially ioized system can be split into bound and free electrons

$$\rho_{\rm e} = \rho_{\rm b} + \rho_{\rm f}$$

- Intermediate scattering function
- $N_{\rm e}F_{\rm ee}^{tot} = \langle \rho_{\rm b}(\vec{k},t)\rho_{\rm b}(-\vec{k},t)\rangle + 2\langle \rho_{\rm f}(\vec{k},t)\rho_{\rm b}(-\vec{k},t)\rangle + \langle \rho_{\rm f}(\vec{k},t)\rho_{\rm f}(-\vec{k},t)\rangle$



Born-Mermin approximation

Fluctuation-dissipation theorem :

$$S_{\rm ee}^0(k,\omega) = -\frac{\epsilon_0 \hbar k^2}{\pi e^2 n_{\rm e}} \frac{{\rm Im}\epsilon^{-1}({\bf k},\omega)}{1 - \exp(-\hbar\omega/k_{\rm B}T_{\rm e})}$$

$$\epsilon^{\text{RPA}}(\vec{k},\omega) = 1 - \frac{1}{\epsilon_0 \Omega_0 k^2} \sum_p e^2 \frac{f_{p+k/2}^e - f_{p-k/2}^e}{\Delta E_{p,k}^e - \hbar(i\omega + i\eta)}$$

• Mermin ansatz :

$$\epsilon_M(k,\omega) = 1 + \frac{\left(1 + \frac{i\nu(\omega)}{\omega}\right)\left[\epsilon^{\text{RPA}}(k,\omega + i\nu(\omega)) - 1\right]}{1 + i\frac{\nu(\omega)}{\omega}\frac{\epsilon^{\text{RPA}}(k,\omega + i\nu(\omega)) - 1}{\epsilon^{\text{RPA}}(k,0) - 1}}$$

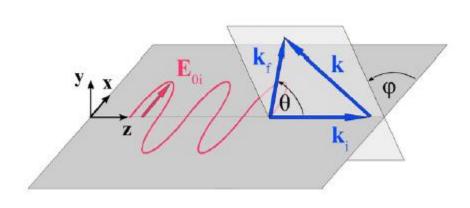
• $\nu(\omega)$ is the dynamic collision frequency via Born approximation.

Glenzer and Redmer, RMP 81, 1625(2009)



44

Back and forward scattering



 The momentum transfer depends on the scattering angle

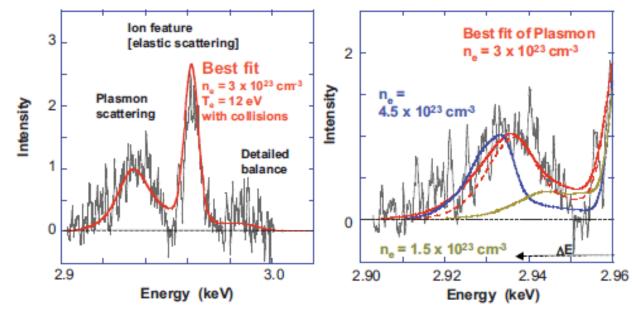
$$k = |k_{\rm f} - k_{\rm i}| = \frac{4\pi}{\lambda_{\rm i}} sin(\theta/2)$$

- Dimensionless scattering parameter $\alpha = \frac{1}{k\lambda_{\rm sc}} = \frac{l}{2\pi\lambda_{\rm sc}}$
 - *l* is the electron density fluctuation
 - $\lambda_{
 m sc}$ is the screening length
- Collective scattering: ($\alpha > 1$)
 - the scattering reflects the electron density fluctuations
 - Plasmon features
- Non-collective scattering:($\alpha < 1$)
 - the scattering reflects the velocity distribution of electrons
 - Compton features



Glenzer and Redmer, RMP 81, 1625(2009)





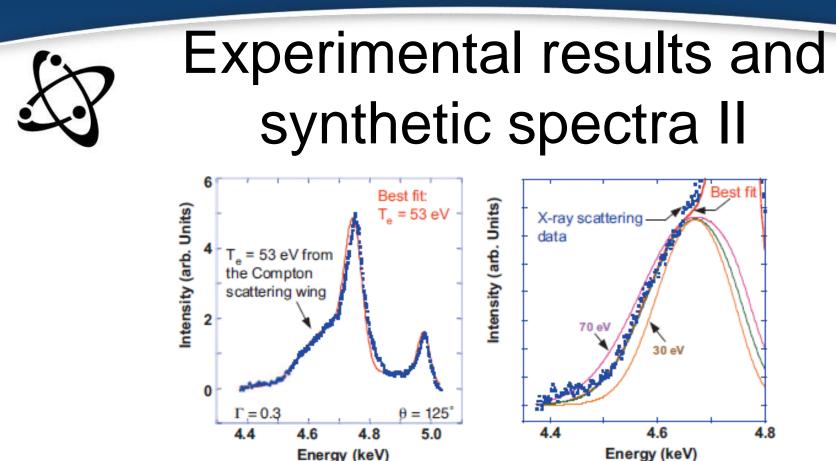
 Forward scattering: collective behavior

Glenzer et al., PRL 98, 065002(2007)

- Dispersion relation determines the electron density
- Detailed balance gives the electron temperature

46





- Back scattering:
 - Compton scattering
 - Non-collective behavior
 - Line width $\,^{\propto}$ Fermi energy

Glenzer et al., PRL 90, 175002(2003)

47





Thanks!







Collisional Radiative Model

•

 T_{j+1}

 S_{j-1}

j+1

1

- Assumptions:
 - Homogeneous plasmas
 - The expansion velocity is constant

$$v_{\text{exp}} = \text{const}$$
 $N(r,t) = N(r',t')(t'/t)^3$

The relative density of charge state is given by

$$n_j = \frac{N_j}{N_{\rm T}}$$

$$\frac{\partial n_j}{\partial t} = N_e S_{j-1} n_{j-1} + N_e T_{j+1} n_{j+1} - N_e (S_j + T_j) n_j \qquad S_j$$
$$\frac{\partial T_e}{\partial t} = \frac{-2T_e}{t} + \frac{2}{3N_e} (P_{\text{recom}} - P_{\text{brem}} - P_{\text{ioniz}}) - P_{\text{ei}} \qquad -\frac{\partial T_i}{\partial t} = \frac{-2T_i}{t} + P_{\text{ei}} \qquad -\frac{\partial T_i}{\partial t} = \frac{-2T_i}{t} + P_{\text{ei}}$$



Collisional Radiative Model

Ionization by collisions

$$S(Z) = 2.43 \times 10^{-6} \xi_Z T_{\rm e}^{3/2} exp(-u)/u^{7/4} \ cm^3 s^{-1}$$

• Recombination: $T_j = \alpha_j + \beta_j + D_j$ Radiative + Three body + Dielectronic

 $\alpha(Z+1) = 5.2 \times 10^{-14} (Z+1) u^{1/2} (0.429 + 0.5 \ln u + 0.469 u^{-1/2}) \ cm^{-3} s^{-1}$

$$D(Z+1) = \frac{1}{T_e^{3/2}} \sum_{j=2}^{Z-1} \sum_{m=1}^{4} C_{mj} \exp(-E_{mj}/T_e)$$

 $\beta(Z+1) = 8.75 \times 10^{-27} N_{\rm e}(Z+1)^3 / T_{\rm e}^{9/2} \ cm^{-3} s^{-1}$





Collisional Radiative Model

Bremsstrahlung-radiation

$$P_{\rm brem} = 9.6 \times 10^{-14} \bar{Z} N_{\rm e}^2 T_{\rm e}^{1/2} \ eV s^{-1} cm^{-3}$$

Ionization energy loss

$$P_{\text{ioniz}} = \sum_{j=1}^{N} nZ - 1N_{\text{e}}N_{j}S_{j}\chi_{Z} \ eVs^{-1}cm^{-3}$$

Energy equilibrium

$$P_{\rm ei} = \frac{T_{\rm e} - T_{\rm i}}{\tau_{\rm eq}} \ eVs^{-1}$$

Three-body recombination

$$P_{\rm recom} = \sum_{j=2}^{nZ} N_{\rm e} N_j \beta_j \chi_Z \ eV \, s^{-1} \, cm^{-3}$$

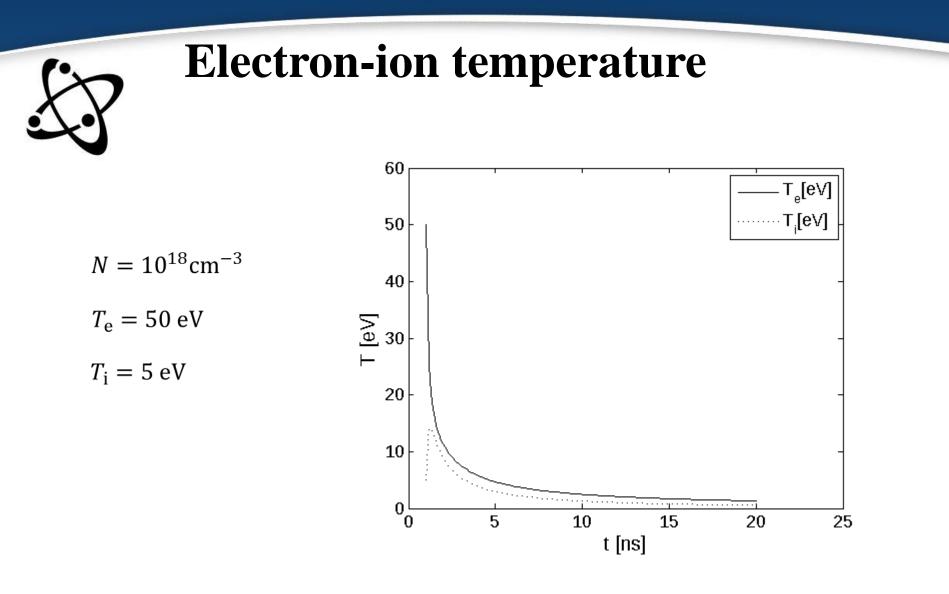
Continuum Lowering



Non-equilibrium Nickel plasmas

- Nickel is a transition element.
- Nickel plays an essential role for the preparation of special stainless-steels and alloys to the production of antimagnetic screens and chemical reactive films.
 - D.B. Chrisey, G.K. Hubler, Pulsed Laser Deposition of thin films, Wiley, 1994
 - S.S. Harilal et al J. Appl. Phys. 114, 203302 (2013)
 - L. Torrisi et al Plasma Phys. Rep. 34, 547 (2008)
 - S. Amoruso et al, Appl. Phys. A. 89, 1017 (2007)

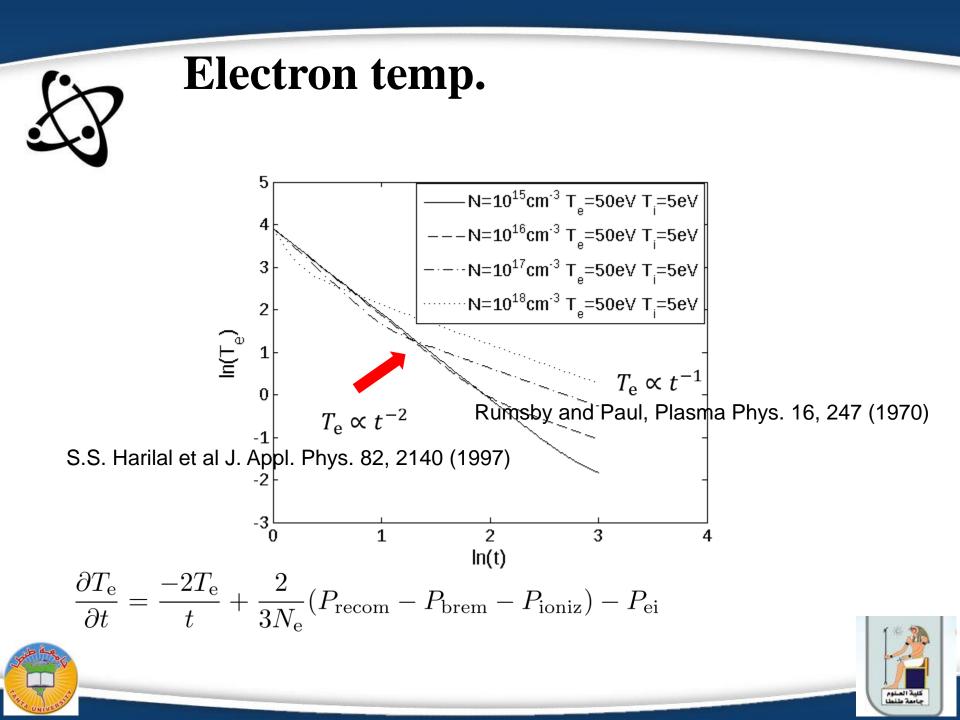






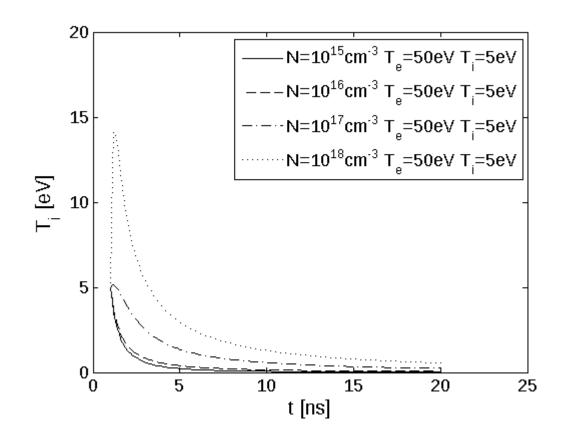
M. Shihab, G.H. Abou-koura, N.M. El-Siragy, Appl. Phys. B (2016) 122-146



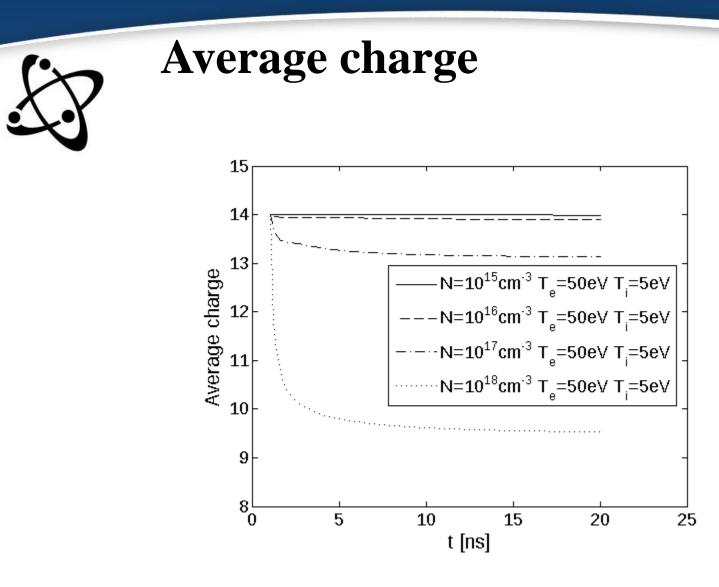




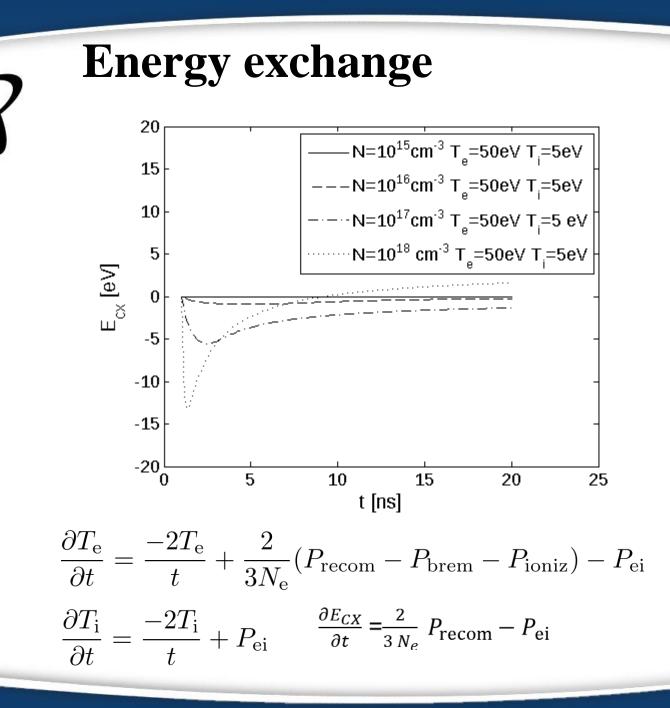
Ion temp.















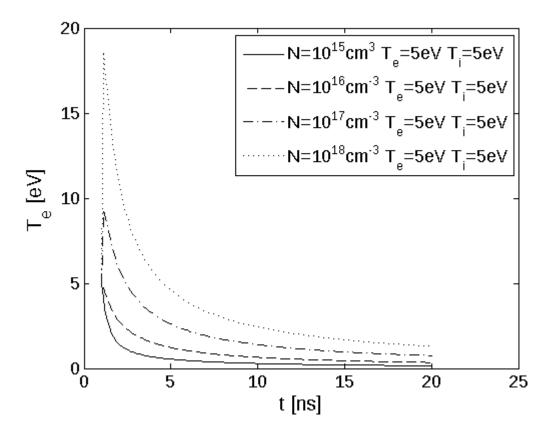
Equilibrium Ni plasmas





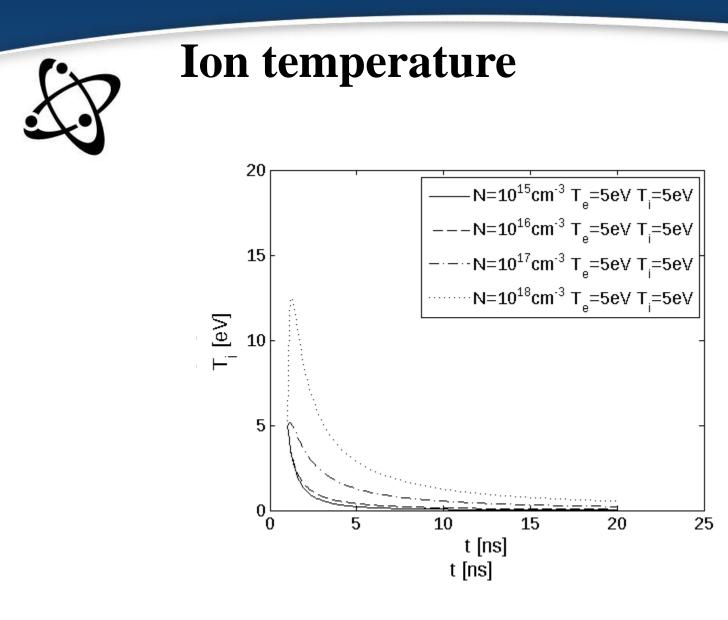


Electron temperature

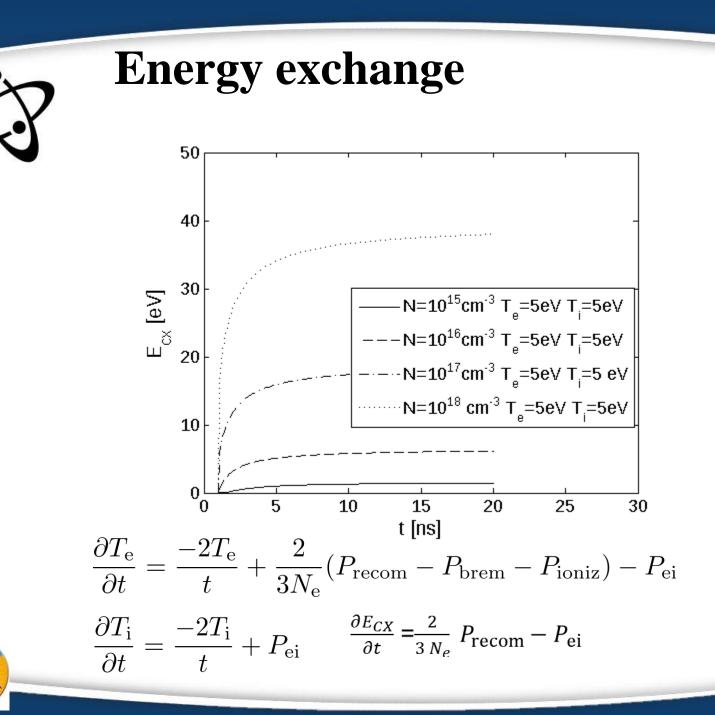




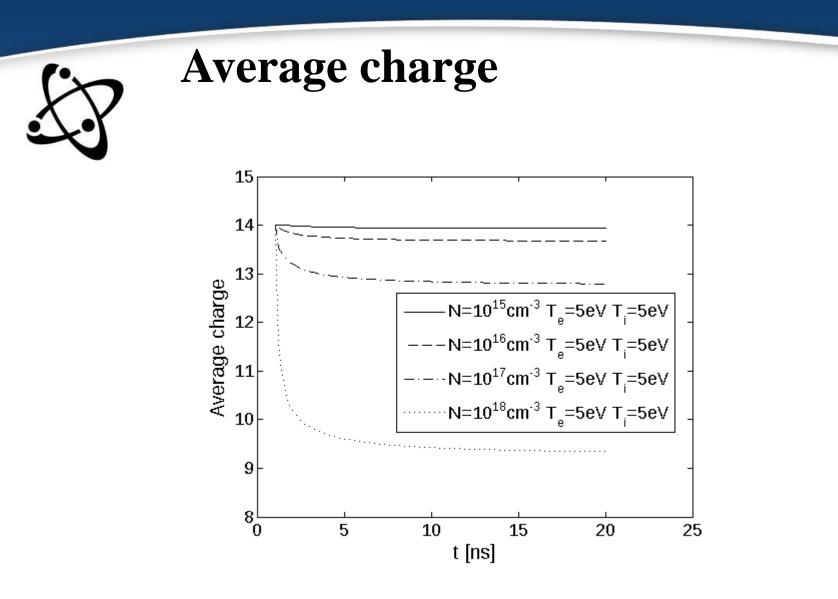
















- The expansion of the plasma after the laser pulses not perfectly adiabatic.
- The three-body recombination acts as an energy source and heats the electrons and consequently changes the electron temperature time dependence.
- Three-body recombination leads to a decrease and freezing of the average charge of the plasma.
- Three-body recombination prolongs the electron-ion relaxation time.
- OUTLOOK: Carbon and Aluminum plasmas. Non-homogeneous expansion





Thanks!

GEFÖRDERT VOM



Bundesministerium für Bildung und Forschung



Ronald Redmer





