

Quantum Plasma

د/ إبراهيم القماش
علم - الفيزياء

* Contents:

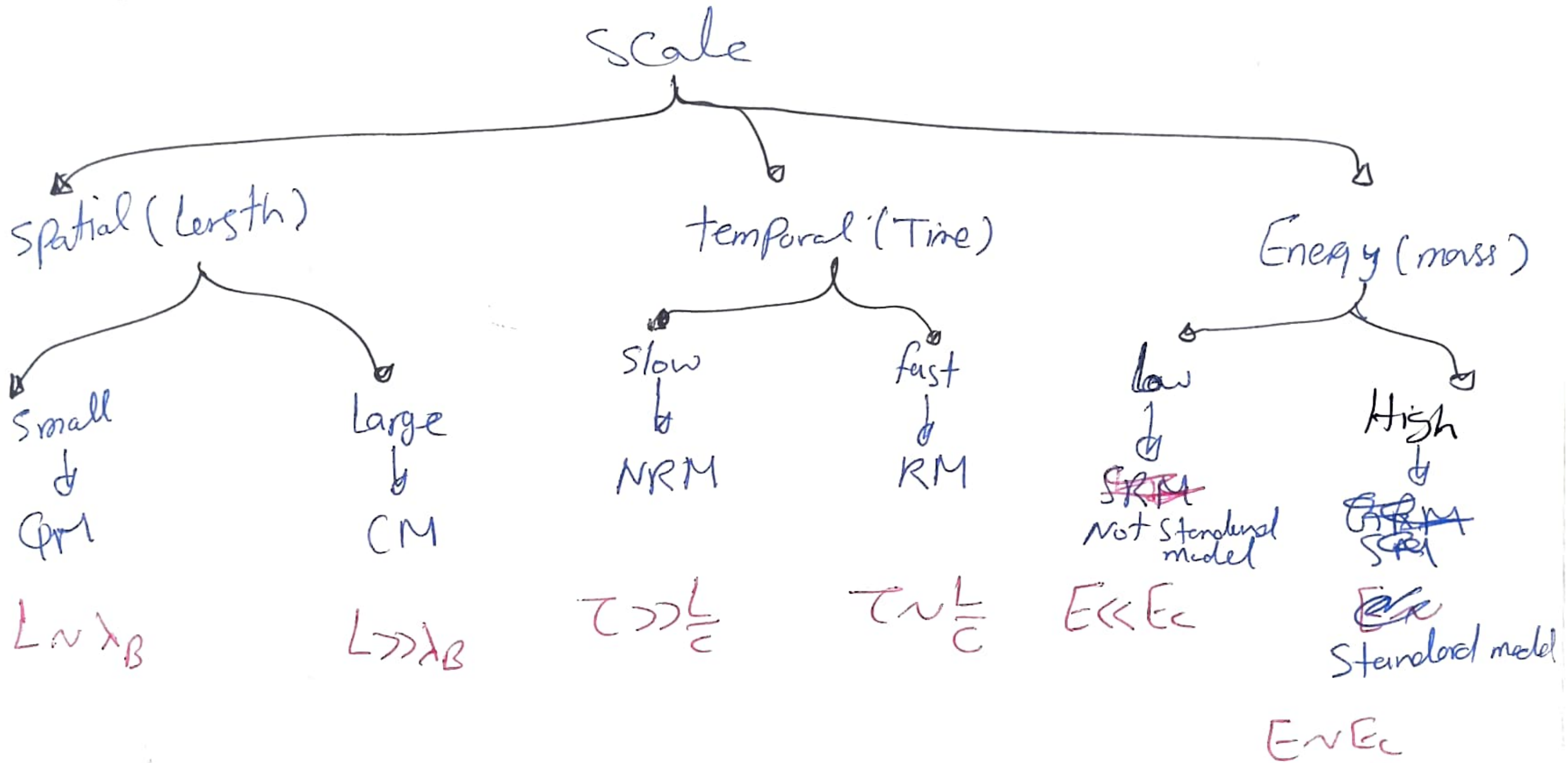
- i - Scale: Definition & types.
- ii - Principles & concepts: classical Vs Quantum
- iii - Quantum Plasma: Definition & Parameters & Existence.
- iv - plasma Models: classical Vs Quantum
- v - Quantum effects:

1- Scale:

• Definition: the size or scope at which physical phenomena are observed or analyzed.

• Fundamental Quantities: Length (L) & Time (T) & Mass (M)

• Types:



$$E_c \approx \frac{mc^3}{eh} \sim 10^{18} \text{ V/m}$$

↳ Schwinger limit

QED effects

Pair Production

concepts & principles: Classical vs Quantum

BC / Aspects		Classical	Quantum
1	Nature	Deterministic	Probabilistic
2	Identity	Distinct	Neither & Duality
3	Trajectory	Definite	Jumping
4	Space	Physical	Hilbert
5	State representation	$\vec{r}, \vec{p}$	$\psi(\vec{r}, t)$
6	State evolution	Newton	Schrodinger
7	Superposition	Discrete state	Mixed state
8	observables	measured values Continuous	Operator Quantized / Discrete
9	observer	Passive	Active
10	Measurements	Exact / Simultaneous / Certain	uncertain Approximate / ordered / uncertain.
11	Statistics	Maxwell-Boltzmann	Fermi-Dirac Bose-Einstein
12	Causality	Absolute / well-defined	Relative / ill-defined
13	Exclusion	Not Applicable	Applicable
14	Entanglement / Locality	Local	Nonlocal
15	tunneling	Not possible	Possible

III → Quantum Plasma:

* Definition: ~~Classical~~ classical plasma + Quantum effects (For it can be called "Semi-classical Plasma").

* Parameters:

	<u>Classical</u>	<u>Quantum</u>
Length ^{scale} :	L	λ_D (Wigner-Seitz radius)
Time:	ω_p	ω_p
Speed:	v_{th}	v_F
Temperature:	T	T_F
Energy:	E_{th}	E_F
Shielding:	λ_D (Debye)	λ_F (Fermi)
Collision:	ν_e (kT)	ν_F (kT_F)

* Existence conditions:

	<u>Classical</u>	<u>Quantum</u>
①	$L \gg \lambda_D$	$r_{ws} \gg \lambda_D$
②	$n \lambda_D \gg 1$	$n \lambda_D \gg 1$
③	$v_e \ll \omega_p$	$v_F \ll \omega_p$

* Conclusion:

- ① $n > n_Q = \left(\frac{2\pi m k_B T}{h^2} \right) \sim 10^{29} \text{ cm}^{-3}$
 Gold. $n \sim 10^{22} \text{ cm}^{-3}$
 Exp. $T_F \sim 10^4 \text{ K}$
- ② $T_F \gg T$
- ③ Degeneracy $\xi = \frac{P_F}{mc^2} \gg 1$

* Classical Plasma: low density + Higher temperature + MB statistics

* Quantum Plasma: High n + Low T + FD statistics

Plasma Models: classical vs Quantum

Note that: the quantization appears only in particles not fields (EPR)

Model	Classical	Quantum
i - PM	<u>Newton - Maxwell</u> $\frac{d\vec{p}_i}{dt} = \sum_j \vec{F}_{ij}$ + Maxwell	<u>Schrodinger - Maxwell</u> $\hat{H} \psi(\vec{r}, t) = H \psi(\vec{r}, t)$ + Maxwell ;
ii - KM	<u>Vlasov - Maxwell</u> $\frac{\partial f}{\partial t} + \vec{v} \cdot \vec{\nabla} f + \frac{\vec{E}}{m} \cdot \vec{\nabla}_v f = 0$ $f \geq 0$ + Maxwell Moments	<u>Wigner - Maxwell</u> $\frac{\partial f}{\partial t} + \frac{\vec{p}}{m} \cdot \vec{\nabla} f + \frac{i}{\hbar} [\hat{H}, f] = 0$ f : Wigner DF $f \geq 0$ + Maxwell Moments
iii - FM	<u>Euler - Maxwell</u> Continuity momentum + Maxwell Moments	<u>Euler - Maxwell</u> Continuity momentum + Quantum effects + Maxwell Moments
	ii - Moments $\rightarrow$ iii	i - Madelung transfer $\rightarrow$ iii ii - Moments $\rightarrow$ iii

Madelung transformation

V- Quantum effects:

① Bohm potential:

- It is a quantum hydrodynamic effect.
- It includes: Wave-particle duality & Quantum tunneling & Quantum diffraction & Quantum Diffusion.
- It emerges naturally when transferring Schrodinger to fluid.
- It arises due to wave function nature (spreading)
- Bohm potential: $V_B = \frac{\hbar^2}{2m} \frac{\nabla^2 \sqrt{n}}{\sqrt{n}}$

$$\hbar \rightarrow 0, V_B \rightarrow 0$$

$$m \rightarrow \infty, V_B \rightarrow 0$$

$$\frac{1}{\sqrt{n}} \rightarrow \infty, V_B \rightarrow 0$$



- Schrodinger eq:

$$i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \nabla^2 \psi + V \psi$$

- Madelung transformation: $\psi = \sqrt{n} e^{iS/\hbar}$

$$\text{def: } \psi(\vec{r}, t) = \sqrt{n(\vec{r}, t)} e^{iS(\vec{r}, t)/\hbar}$$

$$\vec{\nabla} S = \vec{p}$$

$$n = |\psi|^2$$

- the Quantum Hydrodynamic:

$$\frac{\partial n}{\partial t} + \nabla \cdot (n \vec{u}) = 0$$

$$m n \left[\frac{\partial \vec{u}}{\partial t} + \vec{u} \cdot \nabla \vec{u} \right] = -\nabla V_C \neq \nabla V_B$$

$$\vec{F}_B = -\vec{\nabla} V_B$$

↳ Bohm force.

⑥

Quantum tunneling: the total Energy

$$E = V_C + V_B$$

So, even if the classical energy of particle is small compared to the potential, the Bohm potential increases the potential of the particle and help him tunnel.

- Quantum diffraction / Diffusion: The particle spreads up when it passes through small slit instead of moving in a straight line.
due to ∇^2 operator.

- Wave-particle duality: the number density is connected with the wave nature
 $n = |\psi|^2$.

② Degeneracy pressure

- Due to Pauli exclusion principle: when the system ~~become~~ compressed and become so dense, this pressure arises to prevent more compression, because fermions can't occupy the same quantum state.
- It is discovered in astrophysics objects.
- There are 3 possibilities of implosion: $\left(\begin{array}{l} \text{Gravity} > \text{Thermal} \\ \text{ordinary matter} > \text{Nuclear force} \end{array} \right)$
 - IF $M_{\text{star}} \leq 1.4 M_{\odot}$: light star $\rightarrow$ White Dwarfs stars
 $\rightarrow$ Chandrasekhar limit ~~stable~~ stable.
 - $M_{\text{star}} > 1.4 M_{\odot}$: white dwarfs $\rightarrow$ stable Neutron stars
 $\rightarrow$ neutrons.
 - $M_{\text{star}} > 2.5 M_{\odot}$: Neutron stars $\rightarrow$ stable Black hole
 $\rightarrow$ Tolman-Oppenheimer Volkoff limit (TOV).
- Chandrasekhar equation of State:

$$P = A f(\xi)$$

$$A = \frac{\pi m^4 c^5}{3 \hbar^3}$$

$$f(\xi) = \left[\xi(2\xi^2 - 3) \sqrt{1 + \xi^2} + 3 \sinh^{-1} \xi \right]$$

Degeneracy parameter $\xi = \frac{P_F}{m c} = \frac{V_F}{c} = \left(\frac{n}{n_Q} \right)^{1/3}$; $n_Q \sim 10^{29} \text{ cm}^{-3}$

$= \frac{E_F}{\hbar \omega_p} = \frac{\text{Fermi energy}}{\text{Plasma energy}}$

$P_F \begin{cases} \xi \ll 1: \text{NR} : \propto n^{5/3} \\ \xi \gg 1: \text{UR} : \propto n^{4/3} \end{cases}$ $E_F \gg E_p$

* Note: $\begin{cases} \text{Explosion: thermal} > \text{gravity} \\ \text{Implosion: thermal} < \text{gravity} \end{cases}$

Correlation - Exchange :

Electrons don't interact through Coulomb potential; but through.

(*) Exchange: - Due to Pauli exclusion principle, identical electrons can't occupy the same quantum state.

- When the electrons exchange OR swapping its position with each other leads to

$$\psi(\vec{r}_1, \vec{r}_2) = -\psi(\vec{r}_2, \vec{r}_1)$$

- Parallel Electrons with parallel spin don't like to stay close to each other, so they repel each other $\rightarrow$ Repulsion force.

(*) Correlation: - Due to spatial distribution nature of the electrons, the electrons move correlated to each other.

- When the electrons move, they adjust its position to become close to each other $\rightarrow$ Attractive Force.

$$F_{xc} = -\nabla(V_x + V_c)$$

$$V_{xc} = a_1 n^{\frac{1}{3}} + a_2 n^{\frac{2}{3}}.$$